It is easy to criticize this book on the basis of its lack of appeal to the non-technical reader; however, such appeal was hardly the author's intent. One is sure to find no practical suggestions here about how to observe variable stars. Nevertheless the book will undoubtedly become a classic in the literature on the subject. It should be in every variable star observer's library whether he is professional, amateur, technical, non-technical, or otherwise.

Clinton B. Ford Wilton, Connecticut

ANNOUNCEMENT: Revised Sequence for SS Cygni.

A number of observers have noted apparent discrepancies in the sequence for 213843 SS Cygni. Wayne Lowder discovered in the literature a photoelectrically measured sequence for the star (Lenouvel & Daguillon 1956, <u>Journal des Obs. 39</u>, 9). This sequence differs in some respects from that shown on the AAVSO charts. Three Stamford Observatory plates were measured on the Yale University Observatory's microdensitometer by Charles Scovil, confirming the work of Lenouvel and adding three new comparison stars convenient to the variable. The new sequence reproduced below combines this work, and also confirms the changes suggested by the AAVSO visual observers.

The new sequence should be put into use by observers on April 1, 1974 (and not before, for uniformity). Copies of revised charts including a new "e" by Clinton Ford may be obtained from AAVSO Headquarters.

The chart is reproduced to 'd' scale (20"=1mm). Underlined magnitudes are photoelectric (visual), all others are photovisual. South is up.

